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Feb. 10th, 2015

THE INCOMPRESSIBLE EULER LIMIT

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Outline

Two expansions

Hydrodynamic limits

The Vlasov equation

The incompressible Euler limit

THE INCOMPRESSIBLE EULER LIMIT

Lecture 3

The Boltzmann equation and Vlasov equation

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A singular perturbation problem

Note that small Knudsen limit $Kn \rightarrow 0$ to the Boltzmann equation corresponds to a singular perturbation problem:

$$Kn(\partial_t F + \xi \cdot \nabla_x F) = Q(F, F).$$

Thus, formally, as long as $\partial_t F + \xi \cdot \nabla_x F$ is uniformly bounded in the zero Knudsen limit, we may argue that

$$Q(F,F)
ightarrow 0$$
, as $Kn
ightarrow 0$

In other words, as $Kn \rightarrow 0$,

 $F \rightarrow M$ in suitable sense.

Thus, zero Knudsen limit, we may hope that the Boltzmann equation behaves like the Euler equations.

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Question: Can this singular perturbation limit be justified rigorously ?

The compressible Euler limit

Consider the Boltzmann equation $\varepsilon := Kn$:

$$\varepsilon \Big(\partial_t F + \xi \cdot \nabla_x F \Big) = Q(F, F).$$

• The Hilbert expansion D. Hilbert, Begrndung der kinetischen Gastheorie, Math. Ann. 72 (1912), 562-577.

Expand *F* as a formal power series of $\varepsilon = Kn$:

$$F = \sum_{n=0}^{\infty} \varepsilon^n F_n = F_0 + \varepsilon F_1 + \varepsilon^2 F_2 + \cdots$$

◊ L.H.S.:

$$\varepsilon({}_{t}F_{0} + \xi \cdot \nabla_{x}F_{0}) + \varepsilon^{2}({}_{t}F_{1} + \xi \cdot \nabla_{x}F_{1}) + \varepsilon^{3}({}_{t}F_{2} + \xi \cdot \nabla_{x}F_{2}) + \cdots$$

\$\$\$ \& R.H.S.:

$$Q(F_0, F_0) + 2\varepsilon Q(F_1, F_0) \\ + \varepsilon^2 (2Q(F_2, F_0) + Q(F_1, F_1)) + \cdots$$

Compare various orders in ε :

$$\begin{array}{ll} \mathcal{O}(1): & Q(F_0, F_0) = 0 \implies F_0 = M. \\ \varepsilon: & 2Q(F_1, F_0) = \partial_t F_0 + \xi \cdot \nabla_x F_0 =: \mathcal{S}_0, \\ \varepsilon^2: & 2Q(F_2, F_0) = \partial_t F_1 + \xi \cdot \nabla_x F_1 - Q(F_1, F_1) =: \mathcal{S}_1, \end{array}$$

For the solvability of $L_M(F_1) := 2Q(F_1, F_0) = S_0$, by the Fredholm alternative,

 \mathcal{S}_0 is orthogonal to the kernel of L_M^*

But $\text{Ker}L_M^*$ is spanned by the collision invariants,

 $\langle S_0, \psi_{\alpha} \rangle = 0, \quad \alpha = 0, 1, \cdots, 4.$: Euler equations.

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- Some relevant references.
 - *I.* Russel E. Caflisch: The fluid dynamic limit of the nonlinear Boltzmann equation. CPAM. 33 (1980), no. 5, 651D666.
 - 2. Shih-Hsien Yu: Hydrodynamic limits with Shock Waves of the Boltzmann Equation. CPAM 58 (2005), 409-443.
 - 3. Laure Saint-Raymond: A mathematical PDE perspective on the Chapman-Enskog expansion. Bulletin of AMS, 51 (2014), 247-275.

cf. 1. The compressible Navier-Stokes limit: The Chapman-Enskog expansion

2. The incompressible Euler limit: $M \rightarrow 0$.

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- 1. Earle H. Kennard: Kinetic theory of gases (1938).
- 2. Stephen G. Brush: The kinetic theory of gases: An anthology of classic papers with historical commentary. Imperial College Press, 2003.

The Vlasov equation



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If you first hear of the Vlasov equation, you might ask the following questions:

- QA 1: What is the Vlasov equation ?
- QA 2: When do we use the Vlasov equation ?
- QA 3: What are the relations with other fluid equations ?

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What is the Vlasov equation ?

The Vlasov equation = collisionless Boltzmann equation

Physical situation: Consider an ensemble of particles moving in a mean-field force fields (e.g. a many-body particle systems in mean-field setting)

Let $F = F(x, \xi, t)$ be a one-particle distribution function of particles and we assume that the collisions between particles are secondary, in contrast collisions (interactions) between particle and fields are important.

$$\frac{d}{dt}F(x(t),\xi(t),t)=\partial_tF+\dot{x}(t)\cdot\nabla_xF+\dot{\xi}(t)\cdot\nabla_\xi F=0,$$

or equivalently

$$\partial_t F + \frac{\xi}{m} \cdot \nabla_x F + E(x,t) \cdot \nabla_\xi F = 0.$$

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Let V = V(x, y) be the pairwise potential between particles at position *x* and *y*, respectively. In this case, the self-consistent force field

$$E(x,t) = -\nabla_x \iint_{\mathbb{R}^3 \times \mathbb{R}^3} V(x,y)F(y,\xi_*,t)d_*dy$$

= $-\nabla_x \int_{\mathbb{R}^3} V(x,y)\rho(y,t)dy.$

e.g. 1. Electrostatic potential generated by a charge q:

$$V(x,y) = rac{q}{4\pi} rac{1}{|x-y|}$$
 : repulsive

2. Gravitational potential generated by a mass *m*:

$$V(x,y) = -\frac{Gm}{|x-y|}$$
 : attractive

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Thus, the self-consistent Vlasov equation reads as

$$\partial_t F + \frac{\xi}{m} \cdot \nabla_x F + E(x,t) \cdot \nabla_\xi F = 0,$$

$$E(x,t) = -\nabla_x \int_{\mathbb{R}^3} V(x,y) \rho(y,t) dy.$$

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When do we use the Vlasov equation ?

In plasma physics, the equation was first suggested for description of plasma by Anatoly Vlasov in 1938 " A. A. Vlasov (1938). "On Vibration Properties of Electron Gas". J. Exp. Theor. Phys. (in Russian) 8 (3): 291"



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• Plasma physics



Plasma = the fourth fundamental state of gases, completely ionized gases.

e.g., gas inside light bulb.

 $\mathsf{lce} \implies \mathsf{water} \implies \mathsf{vapor} \implies \mathsf{plasma}.$

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The Vlasov-Maxwell system

Vlasov equations for electrons and ions + Maxwell equations for electric and magnetic force fields.

• Dynamic variables

 $F_i = F_i(x, \xi, t), F_e = F_e(x, \xi, t)$: distribution functions for ion and electr E = E(x, t), B = B(x, t): electric and magentic fields density.

We set a relativistic velocity related to momentum ξ :

$$v_{lpha}(\xi) = rac{\xi}{\sqrt{m_{lpha}^2 + |\xi|^2/c^2}}, \quad lpha = i, e.$$

where c is the speed of light. Then it is easy to see that

 $|\mathbf{V}_{\alpha}(\xi)| < \mathbf{C}.$

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The Vlasov-Maxwell system

$$\partial_t F_\alpha + \mathbf{v}_\alpha \cdot \nabla_x F_\alpha + \mathbf{e}_\alpha \left(\mathbf{E} + \frac{\mathbf{v}_\alpha}{\mathbf{c}} \times \mathbf{B} \right) \cdot \nabla_\xi F_\alpha = \mathbf{0}, \\ \partial_t \mathbf{E} = \mathbf{c} \nabla \times \mathbf{B} - \mathbf{j}, \quad \nabla \cdot \mathbf{E} = \rho, \\ \partial_t \mathbf{B} = -\mathbf{c} \nabla \times \mathbf{E}, \quad \nabla \cdot \mathbf{B} = \mathbf{0},$$

where ρ and *j* are the charge and current densities:

$$\rho = C(d) \int \sum_{\alpha} e_{\alpha} F_{\alpha} d\xi, \text{ charge density}$$

$$j = C(d) \int \sum_{\alpha} v_{\alpha} e_{\alpha} F_{\alpha} d\xi \text{ current density.}$$

cf. Small data and global existence:

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A single species plamsa

• Electron gun (laser) and plamsa sheath



Only one species of charged particles, say electrons or ions.

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The Vlasov-Poisson system

Recall a Vlasov-Maxwell system for a single species.

$$\partial_t F + \mathbf{v}(\xi) \cdot \nabla_x F + \mathbf{e} \Big(E + \frac{\mathbf{v}(\xi)}{c} \times B \Big) \cdot \nabla_\xi F = 0,$$

$$\frac{1}{c} \partial_t E = \nabla \times B - \frac{j}{c}, \quad \nabla \cdot E = \rho,$$

$$\frac{1}{c} \partial_t B = -\nabla \times E, \quad \nabla \cdot B = 0.$$

Consider a regime where

$$|\mathbf{v}(\xi)| \ll \mathbf{c}, \quad \partial_t \mathbf{B} \approx \mathbf{0}, \quad \partial_{\mathbf{x}_i} \mathbf{B}_i \approx \mathbf{0}.$$

i.e., letting $c \rightarrow \infty$ and B = 0. Then, the Vlasov-Maxwell system becomes

$$\partial_t F + \mathbf{v}(\xi) \cdot \nabla_x F + \mathbf{e} E \cdot \nabla_\xi F = \mathbf{0}, \\ \nabla \cdot E = \rho, \quad \nabla \times E = \mathbf{0}.$$

We set

$$E = \nabla_x \varphi$$

and obtain the Vlasov-Poisson system:

$$\partial_t F + \xi \cdot \nabla_x F + \nabla_x \varphi \cdot \nabla_\xi F = 0, \quad x, \xi \in \mathbb{R}^3, \ t \in \mathbb{R}, \ \Delta \varphi = \rho, \qquad \rho = \int eFd\xi.$$

cf. Rigorous justification: Degond, Ukai '80 in finite-time interval

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• Astrophysics:

A galaxy is a gravitationally bound system consisting of stars, stellar remnants, an interstellar medium of gas and dust, and dark matter

cf. Size of galaxy: From 10^4 to $10^{14},$ observable universe: $\geq 10^{14}\text{-galaxies}$



James Hopwood Jeans: On the theory of star-streaming and the structure of the universe, Monthly Notices of the Royal Astronomical Society, 76 (1915), 70 -84.

cf. Reinhard Rein



Inner beauties of the V-P system

don't tell me that inner beauty is more important than outer beauty, because in this society it isn't. no one gives you a chance to see if you're beautiful on the inside if you aren't on the outside.



FIND YOUR Inner Beauty

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Inner beauties of the V-P system

• The V-P system is a Hamiltonian system.

Recall that the Vlasov-Poisson (V-P) system reads as

$$\partial_t F + \frac{\xi}{m} \cdot \nabla_x F - \nabla_x \varphi \cdot \nabla_\xi F = 0, \quad x, \xi \in \mathbb{R}^3, \ t \in \mathbb{R}, \\ \Delta \varphi = \rho, \qquad \rho = \int F d\xi.$$

Then, the V-P system is equivalent to the ODE system:

$$\frac{dx}{dt} = \frac{\xi}{m}, \quad \frac{d\xi}{dt} = E = -\nabla_x \varphi.$$

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Define a particle path $[x(s) = x(s; x, v, t), \xi(s) = \xi(s; x, v, t)]$:

$$\frac{dx(s)}{ds} = \frac{\xi(s)}{m}, \quad \frac{d\xi(s)}{ds} = E(x(s), \xi(s), s),$$
$$(x(t), \xi(t)) = (x, \xi).$$

Then, along the particle path, we have

$$F(x(s),\xi(s),s)=F_0(x,\xi), \quad s>0, \ x,\xi\in\mathbb{R}^3.$$

Note that $(x, \xi) \rightarrow (x(s), \xi(s))$ is a measure preserving.

We claim: The above ODE system is a Hamilton's ODEs We set

$$H(x,\xi,t) := \frac{1}{2m} |\xi|^2 + \iint_{\mathbb{R}^3 \times \mathbb{R}^3} V(x,y) f(y,\xi,t) d\xi dy$$
$$= \frac{1}{2m} |\xi|^2 + \underbrace{\int_{\mathbb{R}^3} V(x,y) \rho_F(y,t) dy}_{=:\varphi},$$

where potential energy is mean-field.

♦ Hamilton's ODEs

$$\frac{dx}{dt} = \frac{\partial H}{\partial \xi} = \frac{\xi}{m},$$
$$\frac{d\xi}{dt} = -\frac{\partial H}{\partial x} = -\nabla_x \varphi.$$

• Conservation laws.

Define a particle path $[x(s) = x(s; x, v, t), \xi(s) = \xi(s; x, v, t)]$:

$$\frac{dx(s)}{ds} = \frac{\xi(s)}{m}, \quad \frac{d\xi(s)}{ds} = E(x(s), \xi(s), s),$$
$$(x(t), \xi(t)) = (x, \xi).$$

Then, along the particle path, we have

$$F(x(s),\xi(s),s)=F_0(x,\xi), \quad s>0, \ x,\xi\in R^d.$$

Note that $(x, \xi) \rightarrow (x(s), \xi(s))$ is measure preserving.

• Conservation of *L^p*-norm

$$||F(t)||_{L^p} = ||F_0||_{L^p}, \quad t \ge 0.$$

Balanced laws

Consider a linear Vlasov equation:

$$\partial_t F + \xi \cdot \nabla_x F - \nabla_x \varphi \cdot \nabla_\xi F = \mathbf{0}.$$

Conservtion of mass

As before, we define

$$\rho(\mathbf{x},t) := \int F d\xi, \quad j(\mathbf{x},t) = (\rho u)(\mathbf{x},t) = \int \xi F d\xi.$$

Using the relation

$$\xi \cdot \nabla_{\mathbf{x}} \mathbf{F} = \nabla_{\mathbf{x}} \cdot (\xi \mathbf{F}), \quad \nabla_{\mathbf{x}} \varphi \cdot \nabla_{\xi} \mathbf{F} = \nabla_{\xi} \cdot (\nabla_{\mathbf{x}} \varphi \mathbf{F}).$$

We integrate the Vlasov equation with respect to ξ -variable to obtain the continuity equation (local conservation of mass):

$$\partial_t \rho + \nabla_x \cdot (\rho u) = 0$$
, i.e., $\partial_t \rho + \nabla_x \cdot j = 0$.

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• Balance of momentum

We multiply ξ to the Vlasov equation

$$\partial_t(\xi F) + \nabla_x \cdot \left(\frac{\xi \otimes \xi}{m}F\right) + \nabla_\xi \cdot \left(\xi \otimes \nabla_x \varphi F\right) = -\nabla_x \varphi F$$

and integrate the resulting relation with respect to $\boldsymbol{\xi}$ to obtain

$$\partial_t(\rho u) + \nabla_x \cdot (\rho u \otimes u + P) = -\rho \nabla_x \varphi.$$

Conservation of total energy

Define an energy

$$\mathsf{E}:=\int \frac{|\xi|^2}{2}\mathsf{F}d\xi dx.$$

We multiply $\frac{|\xi|^2}{2}$ to the equation to obtain

$$\partial_t \left(\frac{|\xi|^2}{2}F\right) + \nabla_x \cdot \left(\xi \frac{|\xi|^2}{2}F\right) + \nabla_\xi \cdot \left(-\nabla_x \varphi \frac{|\xi|^2}{2}F\right) \\ + \nabla_x \cdot (\varphi \xi F) - \varphi \nabla_x \cdot (\xi F) = 0.$$

We integrate the above relation with $d\xi dx$ to get

$$\frac{d}{dt}\iint \frac{|\xi|^2}{2}Fd\xi dx - \iint \varphi \nabla_x \cdot (\xi F)d\xi dx = 0.$$

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Note that

$$-\iint \varphi \nabla_x \cdot (\xi F) d\xi dx = -\int \varphi \nabla_x \cdot j dx = \int \varphi \partial_t \rho dx$$
$$= \iint V(|x - y|)\rho(y, t)\partial_t \rho(x, t) dy dx$$
$$= \frac{1}{2} \iint V(|x - y|)\partial_t (\rho(x, t)\rho(y, t)) dy dx$$
$$= \frac{d}{dt} \frac{1}{2} \iint V(|x - y|)\rho(x, t)\rho(y, t) dy dx$$
$$= \frac{d}{dt} \frac{1}{2} \int \varphi(x, t)\rho(x, t) dx.$$

Finally, we have the conservation of total energy:

$$\frac{d}{dt}\Big[\iint \frac{|\xi|^2}{2}Fd\xi dx + \frac{1}{2}\int \varphi(x,t)\rho(x,t)dx\Big] = 0.$$

or equivalently,

$$\frac{d}{dt}\left[\iint \frac{|\xi|^2}{2}Fd\xi dx + \frac{1}{2}\int |E(x,t)|^2 dx\right] = 0.$$

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• Conservation of entropy

Note that

$$\partial_t(F \ln F) = (\partial_t F)(1 + \ln F) = -\nabla_x \cdot (\xi F \ln F) + \nabla_\xi \cdot (\nabla_x \varphi F \ln F).$$

We integrate the above relation with respect to $d\xi dx$ to get

$$\frac{d}{dt}\iint F\ln Fd\xi dx=0.$$

The V-P system is a conservative system.

Incompressible Euler limit

• Quasi-neutral limit Let $F = F(x, \xi, t)$ be a kinetic density for electron, and assume that the ion density is constant, say 1, i.e.,

$$\partial_t F + \xi \cdot \nabla_x F - \nabla_x \varphi \cdot \nabla_\xi F = 0,$$

 $\varepsilon \Delta \varphi = 1 - \int F d\xi, \quad \varepsilon : \text{ Debye length.}$

By previous argument, we have local and global balanced laws.

$$\begin{split} \partial_t \int F d\xi + \nabla \cdot \int \xi F d\xi &= 0, \\ \partial_t \int \xi F d\xi + \nabla \cdot \int \xi \otimes \xi F d\xi + \nabla \varphi \\ &= \varepsilon \nabla \cdot (\nabla \varphi \otimes \nabla \varphi) - \frac{\varepsilon}{2} \nabla (|\nabla \varphi|^2), \\ \frac{d}{dt} \Big[\iint \frac{1}{2} |\xi|^2 F d\xi dx + \int \frac{\varepsilon}{2} |\nabla \varphi|^2 dx \Big] = 0. \end{split}$$

cf. E. Grenier, Y. Brenier

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Quasi-neutral limit

We take $\varepsilon \rightarrow 0$ (quasi-neutral limit)

$$\int Fd\xi = 1.$$

We again recall

$$\rho(\mathbf{x},t) = \int F d\xi = 1, \quad J(\mathbf{x},t) = \int \xi F d\xi.$$

and take an ansatz (for perfectly cold electrons)

$$F(x,\xi,t) = \delta(\xi - J(x,t)).$$

to get the incompressible Euler equation:

$$\nabla \cdot \boldsymbol{J} = \boldsymbol{0}, \quad \partial_t \boldsymbol{J} + \nabla \cdot \boldsymbol{J} \otimes \boldsymbol{J} + \nabla \varphi = \boldsymbol{0}.$$

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Summary of Lecture 3

- The Vlasov-Poisson system is a conservative system (mass, energy)
- The compressible, incompressible fluid equations can be formally derived from the Boltzmann and Vlasov-Poisson systems.

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From "The Biological Physicist" August 2005



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- **Bahar**: What advice would you have for a scientist just beginning a career in interdisciplinary science ?
- Strogatz: First,

Second, don't be afraid to work in a completely unfamiliar subject. You can come up to speed amazingly quickly if you have a collaborator in that field, and if you hang around his or her lab for a few weeks. And keep in mind that you bring many advantages as an outsider. You have a different set of tools. You will ask unusual questions. And you don't know know what's impossible.